







# Nuclear Spin Conversion in Gas Phase Recent developments

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# Nuclear spin

#### Electron spin : fine structure of hydrogen atom and Stern-Gerlach experiment

G.E. Uhlenbeck and S. Goudsmit, Naturwissenschaften 47 (1925) 953.

#### Unexplained experimental data for H<sub>2</sub> molecule :



#### Intensity alternation

**Anormal Rotational Specific Heat** 

By similarity with the interpretation of the Helium spectrum involving electron spin, Dennison proposed an interpretation of the abnormal specific heat of H<sub>2</sub> introducing the proton nuclear spin.

David M. Dennison Am. J. Phys. 42, 1055 (1974)



# Nuclear spin isomers

Normal hydrogen H<sub>2</sub> is a mixture of two different types of molecules that have different thermal and optical properties



All molecules having identical non zero nuclear spin atoms (H, F ...) in symmetrical positions exist as nuclear spin isomers and are distinguished with the total nuclear spin of the equivalent atoms.



# Role of equivalent atoms

**Pauli's Principle : the total wavefunction has to be symmetric or antisymmetric with respect to permutation of identical atoms** 

The particle type determines the properties of the total wavefunction

Fermi-Dirac

 $(12)\psi = -\psi$ (12)  $\psi = \psi$ fermions (*I* half-integer) bosons (*I* integer) <sup>1</sup>/<sub>2</sub> : H, <sup>13</sup>C, <sup>15</sup>N, F **0** : <sup>12</sup>**C**, <sup>16/18</sup>**O 1** : D, <sup>14</sup>N 3/2 : CI, Br, Na 5/2 : I, <sup>17</sup>O, AI 2: n/a : <sup>73</sup>Ge, <sup>83</sup>Kr 6:V

**Bose-Einstein** 

Molecular symmetry group  $(12) \Psi^{tot} = \pm \Psi^{tot}$   $\Psi^{tot} = \Psi^{e} \Psi^{vib} \Psi^{rot} \Psi^{ns} = \Psi^{evr} \Psi^{ns}$ 

Molecular symmetry group is used to label the energy levels and Pauli's principle allows only certain combinations.

The total nuclear spin I is connected to the rotational quantum numbers



Each spin isomer can be identified by its own rotation-vibration spectrum

# Spin conversion definition

Spin conversion is a <u>dynamical property</u> of molecules embedded in an environnement and results in the possible change of the total nuclear spin of the molecule (ortho  $\leftrightarrow$  para).

The simplest situation is a gas sample where a molecule undergoes collisions



Such a hidden characteristic is quite robust versus collisions and fields : isomers can be regarded as different species.

# **Physical Origin of Conversion**

Flipping the spin requires a gradient of magnetic field strong enough at the molecular scale

**External magnetic field** 

Collisions with paramagnetic molecules (O<sub>2</sub>) H<sub>2</sub>, H<sub>2</sub>O in rare gas matrices, CH<sub>3</sub>F

**Magnetic field inside the molecule** 

Interaction between spins (spin-spin) - Rotation of charges (spin-rotation) CH<sub>3</sub>F, H<sub>2</sub>CO, C<sub>2</sub>H<sub>4</sub>



Spin-spin



Spin-rotation

If the existence of spin isomers is accepted, the <u>dynamics of conversion</u> of one isomer to the other is not well understood.

#### **Our PROJECT**

Provide experimental observations leading to a quantitative explanation of the mechanism of nuclear spin conversion.

(validation and universality of the Quantum Relaxation Model)

Extract informations about physical properties of molecules and their behavior in their environment.

(collisional efficiency ; surface effect)

Make some predictions through theoretical model and spectroscopic data to evaluate spin conversion rates

(ISM)



« Observation of isomerization rates would require the preparation of a non equilibrium isomeric mixture »

Curl et al J. Chem. Phys. 46(8), 3220 (1967)

One must be able to create a disequilibrium between the populations of the different spin configurations to measure the kinetics of the recovering of the Bolztmann distribution.



**Difficulty : identical physical and chemical properties** 



# No general process for enrichment

 $H_2$  (δω # 170.6 K) : cooling below 20.4 K (boiling temperature)

Selective adsorption : H<sub>2</sub>O

Reaction from enriched reactants : H<sub>3</sub>+

Selective UV laser photodissociation : H<sub>2</sub>CO

Rapid cooling (in matrices) : CH<sub>4</sub> (para - H<sub>2</sub>) ; H<sub>2</sub>O (Ar)

Light-induced drift Clean environment – gas phase



# Light-Induced Drift

#### Vibrational and velocity-selective excitation ...



#### ... produces a change of collisional cross section of one isomer species

R(4,K) absorption lines of of the  $v_3$  band of  ${}^{13}CH_3F$ 

> P(32) line of CO<sub>2</sub> laser (9.6 μm band)





# Record of the enrichment phase followed by the conversion phase in <sup>13</sup>CH<sub>3</sub>F







# "Quantum Relaxation" model

For nonlinear molecules with identical nuclei an important pathway leading to equilibration of nuclear spin statistics isomers is provided by wavefunction mixing induced by the spin-rotation interaction and in some cases the spin-spin interaction.

[...] The spin-rotation mixing may be very important if there is an accidental near degeneracy of the right sort. Then most of the isomerization « funnels » through the near – degenerate levels.

R. F Curl et al J. Chem. Phys. 46(8), 3220 (1967)

Energy levels and transitions for a molecule having <u>no</u> equivalent atoms



Energy levels and transitions for a molecule existing as two isomers





Spin-spin interaction  $\Delta J \leq 2, \ \Delta K \leq 2$ 

# Spin-rotation interaction $\Delta J \leq 1, \ \Delta K \leq 2$



# "Quantum Relaxation" model

# **Ingredients:**

- Energetically close ortho and para levels
- Intramolecular interaction coupling these levels
- Collisional relaxation (Maxwell Boltzmann distribution)



# Verification of the model

#### **Pressure dependence of the conversion rate**

$$\gamma = 2 \left[ V_{\alpha\alpha}, \right]^{2} \underbrace{\Gamma_{\alpha\alpha'}}_{P_{\alpha\alpha'}} (W_{\alpha} + W_{\alpha'}) \underbrace{\Gamma_{\alpha\alpha'}^{2} + \omega_{\alpha\alpha'}^{2}}_{P_{\alpha\alpha'}} (W_{\alpha'} + W_{\alpha'}) \underbrace{\Gamma_{\alpha\alpha'}^{2} + \omega_{\alpha\alpha'}^{2} + \omega_{\alpha\alpha'}^{2$$

Low

# Nuclear spin conversion in ethylene H<sub>2</sub>C=CH<sub>2</sub>





### $H_2^{13}C=CH_2$ $\gamma/P = 5 \ 10^{-4} \ s^{-1}/torr$

Chem. Phys. Lett. 322, 424 (2000)



 $H_2C=CH_2$  $\gamma/P = 5.5 \ 10^{-4} \ s^{-1}/torr$ 

Science. 310, 1938 (2005)

### Verification of the model

#### Temperature dependence of the conversion rate

$$\gamma = 2 |\mathbf{V}_{\alpha\alpha'}|^2 - \frac{\Gamma_{\alpha\alpha'}}{\Gamma_{\alpha\alpha'}^2 + \omega_{\alpha\alpha'}^2} (\mathbf{W}_{\alpha} + \mathbf{W}_{\alpha'})$$



### Verification of the model

#### Level crossing resonance

$$\gamma = 2 |\mathbf{V}_{\alpha\alpha},|^2 \underline{\Gamma_{\alpha\alpha'}}_{\Gamma_{\alpha\alpha}}, (\mathbf{W}_{\alpha} + \mathbf{W}_{\alpha'})$$
$$\Gamma_{\alpha\alpha}^2, + \omega_{\alpha\alpha'}^2,$$

 $F = 1017 V.cm^{-1}$ 







Time sequence with and without Stark field  $T_{tot}=T_0 + T_{St}$   $\gamma_{exp} T_{tot} = \gamma_0 T_0 + \gamma_{St} T_{St}$ Enrichment probe during  $T_0$  (field off)





P. Cacciani, J. Cosléou, F. Herlemont, M. Khelkhal & J. Lecointre, Phys. Rev. A, 69, 032704 (2004)

# Magnetic interactions

	ntera	ction s	spin – spir	ו	C. Puzzarini J. Cosléou, P. Cacciani,			
		Ab init	<i>io</i> experimental		F. Herlemont & M. Khelkhal, Chem. Phys. Lett. 401 357-362 (2005)			
<sup>13</sup> CH <sub>3</sub> F	T <sub>22</sub>	69.2	67.9	3 (25)				
Interaction spin – rotation								
CASSCF, ACES2					Ab initio	experimental		
			$^{12}\text{CH}_3\text{F}$			Klemperer (1971)		
			$C(^{19}F)$	$C_{xx} = C_{yy}$	2.07	4.0 (19)		
$C_{\perp} = (C_{xx} + C_{yy})/2$ $\Delta C_{\perp} = (C_{xx} - C_{yy})/2$				C <sub>zz</sub>	- 52.85	-51.1 (13)		
			$C(H_1)$	$\mathrm{C}_{\perp}$	15.92	14.66 (70)		
				C <sub>zz</sub>	0.67	0.8 (15)		
			<sup>13</sup> CH <sub>3</sub> F			This work		
		<b>C</b> ( <b>H</b> <sub>1</sub> )	$\Delta \mathbf{C}_{\perp}$	- 1.88	1.995 (10)			

#### Consequences for a molecule of C<sub>3v</sub> symmetry

- N being isotrope in the perpendicular plane  $(N_{xx} = N_{yy})$ , the transverse anisotropy is entirely of electronic origin :  $\Delta C_{\perp} = \Delta E_{\perp}$ .
- Non-diagonal components of the spin-rotation tensor are entirely of nuclear origin et appear unscreened by electrons :  $C_{\alpha z} = N_{\alpha z}$  ( $E_{\alpha z} = 0$ )  $\alpha = x,y$

#### Specific selection rules for spin-rotation interaction

•  $\Delta J \le 1$ ,  $\Delta k = 2$ : coupling term proportional to  $|\Delta C_{\perp}|^2$ Exemple:  ${}^{13}CH_3F$ : (J,k) = (21,1) - (20,3)

•  $\Delta J \le 1$ ,  $\Delta k = 1$ : coupling term proportional to  $|C_{\alpha z}|^2$ Exemple:  ${}^{12}CH_3F$ : (J,k) = (28,5) - (27,6)

# CH<sub>3</sub>F is a molecule for which it is possible to derive the complete spin-rotation tensor

(E. Ilisca and K. Bahloul Phys. Rev. <u>A57</u>, 4296 (1998))

# Relaxation rates of coherence $\Gamma$

### **Coherence : created by an** *interaction* **and broken by** *collisions*

b magnetic Int. a Pressure dependence of the nuclear spin conversion rate





# Relaxation rates of coherence $\Gamma$

Analogy with collisional broadening of molecular transitions



Pressure dependence of the nuclear spin conversion rate

Pressure broadening

Despite the final measurements are very different, the treatment of the relaxation developed for the line shape can be transferred to the nuclear spin conversion rate.

### Relaxation rates of coherence $\Gamma$

The relaxation parameter  $\Gamma$ , presented as a phenomenological parameter by Chapovsky, can be efficiently calculated with the help of models developed for collisional broadening of molecular transitions.

Parameter	Fitted Value	Calculated value		
Γ <sub>9,3/11,1</sub>	1.55 10 <sup>8</sup> s <sup>-1</sup> .Torr <sup>-1</sup>	1.59 10 <sup>8</sup> s <sup>-1</sup> .Torr <sup>-1</sup>		
Γ <sub>20,3/21,1</sub>	1.34 10 <sup>8</sup> s <sup>-1</sup> .Torr <sup>-1</sup>	1.21 10 <sup>8</sup> s <sup>-1</sup> .Torr <sup>-1</sup>		
Cacciani, J. Cosléou, F.				
Mol. Struct., <u>780-781</u> ,	Within the quantum relaxation model	Collisional inelastic cross-sections		

P. He C.





1<sup>st</sup> pair: matching of experimental and calculated positions of maxima 2<sup>nd</sup> pair : disagreement !

Rotational dependence of the dipole moment

 $\mu = \mu_0 + \mu_J J (J+1) + \mu_K K^2$ 

 $\mu_{J}$  (exp) = 2.72 (1.71) 10<sup>-5</sup> D (figures in parentheses = 3 times the s.d.)

> $\mu_{\rm J}$  (th) = 1.492 10<sup>-5</sup> D  $\mu_{\rm K}$  (th) = - 3.695 10<sup>-5</sup> D

J. Cosléou, P. Cacciani, F. Herlemont, M. Khelkhal, J. Lecointre & P. Pracna, Phys. Chem. Chem. Phys., <u>6</u>, 352 – 357 (2004)

# **Application to astrophyscics**



Ortho/para ratio (OPR) can be measured from observation of interstellar medium or comets. It allows to derive a spin temperature  $T_{spin}$ 





J. Crovisier, Faraday Discuss., 109, 437 (1998)

# Molecular formation deduced from T<sub>spin</sub>



- Prestellar clouds (L723,...): T<sub>spin</sub>~10 K of H<sub>2</sub>CO close to T<sub>kinetic</sub> indicates the formation of molecules or the thermal equilibration on <u>cold grains</u>
- Others cases (L1498, ...) : an OPR = 3:1 whereas T<sub>kinetic</sub> ~10 K indicates a formation at high temperature before it cooled down

(Dickens, Astrophys. J. 1999)

- For most comets (various distances from the Sun, different origins), the measured *OPR* for  $NH_3$ ,  $H_2O$  et  $CH_4$  correspond <u>always</u> to a spin temperature of about 30 K : no re-equilibration
- In the ice, these molecules could have maintained a memory of their formation temperature in the protosolar nebulae
- « ..., we conclude that the Sun was born in a warm molecular cloud near 30 K, not in a cold dark cloud near 10 K, as is usually assumed. » Kawakita et al, Astrophysical Journal 2005



#### A test of the hypothesis of forbidden conversion in gas phase

Nuclear spin conversion of formaldehyde in star forming regions induced by non reactive collisions using "quantum relaxation model"

In collaboration with C. PUZZARINI (Bologna), S. MARET and C. KAHANE (Grenoble)

Energy differences

Identification of the most important ortho-para pairs

**Spin-rotation interaction V**<sub>op</sub>

non diagonal Terms of the tenso never compared to experiment Ab-initio Calculation of the spin-rotation tensor (method CCSD(T), ACES2 program)

Calculation of wavefunctions (asymmetric top) and interaction terms

Decoherence induced by collisions with H<sub>2</sub>

Value estimated from the broadening of spectral lines  $\Gamma$ = 5 10<sup>7</sup> s<sup>-1</sup> Torr<sup>-1</sup>

Dipole-quadrupole interaction between H<sub>2</sub>CO and H<sub>2</sub> Temperature dependence

#### **4** environments

- Prestellar core (T = 5-10 K,  $n(H_2) = 10^6 \text{ cm}^{-3}$ )
- Protoplanetary disk (30 K, 10<sup>7</sup> cm<sup>-3</sup>)
- Photodissociation region Orion (75 K, 2×10<sup>5</sup> cm<sup>-3</sup>)
- Protostar (100 K, 3×10<sup>8</sup> cm<sup>-3</sup>)





Spin conversion characteristic times are much longer than  $H_2CO$  lifetime (~ 10000 years).

#### Potential sources of nuclear spin conversion :

- proton exchange with interstellar H<sup>+</sup>
- exchange of protons attached to C and O atoms within H<sub>2</sub>COH<sup>+</sup> (protonation of H<sub>2</sub>CO by H<sub>3</sub><sup>+</sup>)
- adsorption-desorption on grains.

# ISM → low pressure range





#### PRESSURE DEPENDENCE OF THE CONVERSION RATE



Eur. Phys. J. D <u>12</u>, 297 (2000)

H<sub>2</sub>CO : observation of an increase of γ at low pressure
(C. Bechtel, E. Elias, B.F. Schramm, J. Mol. Struct., <u>741</u>, 97 (2005) )

#### Pressure reduced to 30 mTorr ; 3 cells of different S/V ratio



# P > 200 mTorr : aligned points

# P < 200 mTorr : acceleration of the conversion



With linear dependence at high pressure substracted.

# Volume $\gamma_g$ and non magnetic surface $\gamma_s$ contributionsNumber of collisionsBinary collisionsSurface collisions



 $N_s + N_g$ 

 $X_s =$ 



 $x_g = 1 - x_s$ 



#### **Conversion rate at lower pressure**



$$p_0$$
 defined by  $N_g = N_s$ 

$$p_0 = \frac{kT}{2\sqrt{2}\sigma} \times \frac{S}{V}$$

Cell type	Cross section ( $A^2$ )	p₀ (mTorr)	$\gamma_g$ (s <sup>-1</sup> /Torr)	γ <sub>s</sub> (s <sup>-1</sup> )
Cylindrical cells 20 cm	4541,53 <i>[3</i> ]	0,31	0,01384(11)	0,30(3)
Cylindrical cells 20 cm	44 <i>[4</i> ]	31,98	0,01453(14)	0,0052(2)
Cylindrical cells 50 cm	4541,53 <i>[3</i> ]	0,4	0,01469(13)	0,135(8)
Stark cells	4541,53 <i>[3</i> ]	0,45	0,01481(22)	0,26(4)
All data	4541,53 [3]	0,4	0,01399(8)	0,151(8)